



Celestial Mechanics

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Introduction

What do we observe?

Epicyles

Kepler's Laws

Two Body Problem

Newton's law of Universal Gravitation

Three body problem

Chaos

N-body systems

Introduction

What do we observe?



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Introduction

What do we observe?



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- ▶ He would have started classifying objects into those that move with respect to the sky background and those that don't.



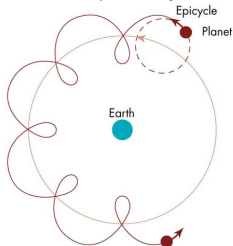
Motion of Mars in the sky.

Introduction

Ptolemy's Model of Epicycles



- ▶ Ptolemy assumed that planets moved around the earth . To explain retrograde motion, he used a complicated system of epicycles (circles around circles) to replicate observations.

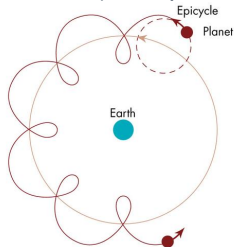


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Ptolemy's Model of Epicycles



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- ▶ The complexity of this model increased as we made more accurate observations.



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- ▶ The orbit of a planet is an ellipse with the Sun at one of the two foci.
- ▶ A line segment joining a planet and the Sun sweeps out equal areas during equal intervals of time.
- ▶ The square of the orbital period of a planet is proportional to the cube of the semi-major axis of its orbit.

Two Body Problem

Setting up the game



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- ▶ These are in all 6 coupled second order non-linear differential equations.

Two Body Problem

Attack



- ▶ We separate the motion into independent components:

$$M \frac{d^2 \vec{R}}{dt^2} = 0$$
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- ▶ The force being radial implies angular momentum conservation. This constraints the motion of two bodies to be in a plane.
- ▶ We go to plane polar co-ordinates and obtain the following equations:

$$\frac{d(\mu r^2 \dot{\theta})}{dt} = 0$$
$$\mu \ddot{r} - \mu \dot{\theta}^2 r = \frac{-Gm_1 m_2}{r^2}$$

Two Body Problem

Finishing Off



- ▶ We define the following variables for convenience: $l = \mu r^2 \dot{\theta}$ and $k = Gm_1 m_2$. We get,

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$$r = \frac{\rho}{1 - e \cos \theta}$$

- ▶ It is easy to see that this is the equation of a conic section. We have solved the two body problem. Cheers!

Three body problem

Failure



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Three body problem

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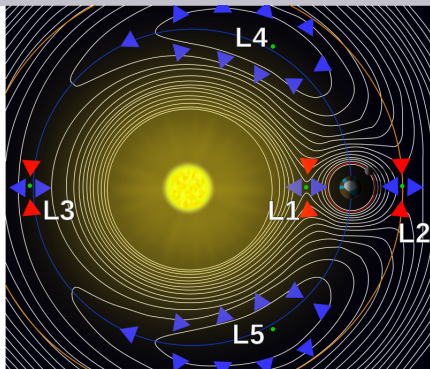
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- ▶ **Is all hope lost then?**

Three Body Problem

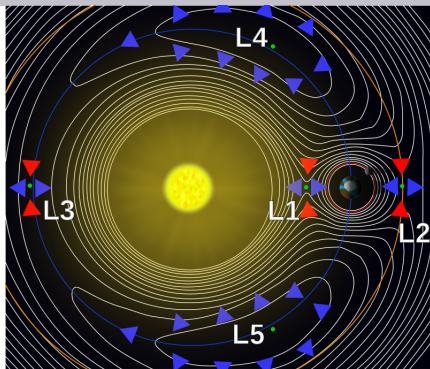
Revival



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Three Body Problem

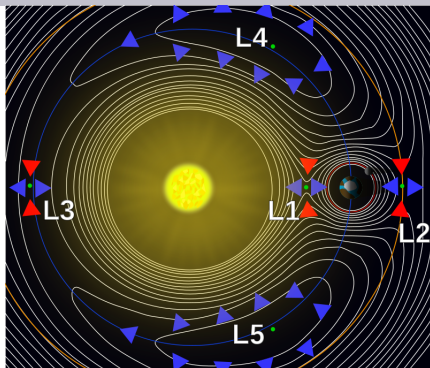
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- ▶ L_2 is useful for placing observational satellites that need absolute absence of sunlight. Examples include WMAP, Planck, etc.
- ▶ A certain type of asteroids known as Trojans are present at L_4 and L_5 of Jupiter.

Chaos

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- ▶ Where does Chaos fit in with celestial mechanics?



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This kind of dichotomy wherein the single case looks bad but the system as a whole behaves nicely is echoed across physics.

- ▶ For example, compare quantum mechanical description of a single electron and a solid object as a whole.

That's all folks!





- ▶ https://en.wikipedia.org/wiki/Two-body_problem
- ▶ <https://en.wikipedia.org/wiki/Chaos>
- ▶ 3 body Sonified -
https://www.youtube.com/watch?v=f_w6JprsXK8
- ▶ Lorenz attractor simulation -
<https://www.youtube.com/watch?v=dP3qAq9RNLg>
- ▶ Simulation of Large Scale structure of the universe -
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